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# Teleseismic Evidence for a Low-Velocity Body Under the Coso Geothermal Area

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Teleseismic *P* wave arrivals were recorded by a dense array of seismograph stations located in the Coso geothermal area, California. The resulting pattern of relative residuals reveals an area showing approximately 0.2-s excess travel time that migrates with changing source azimuth, suggesting that the area is the 'delay shadow' produced by a deep, low-velocity body. Inversion of the relative residual data for three-dimensional velocity structure determines the lateral variations in velocity to a depth of 22.5 km beneath the array. An intense low-velocity body, which coincides with the surface expressions of late Pleistocene rhyolitic volcanism, high heat flow, and hydrothermal activity, is resolved between 5- and 20-km depth. It has maximum velocity contrast of over 8% between 10 and 17.5 km. The shallowest part of this body is centered below the region of highest heat flow; at depth it is elongate in approximately the N-S direction. The hypothesis that this low-velocity body is caused by the presence of partial melt in the middle crust is consistent with the local seismic, geologic, and thermal data.

## INTRODUCTION

Analysis of teleseismic *P* waves recorded above geothermal systems has proven to be an effective method of determining the seismic properties of the underlying crust and upper mantle. A model of the seismic velocity structure provides useful constraints on the compositional and physical states of the geothermal system at depth. Teleseismic *P* wave delays have been used extensively to infer velocity structure at several geothermal systems, including The Geysers, California [Steeles and Iyer, 1976a; Iyer et al., 1979], Yellowstone, Wyoming [Iyer, 1975; Iyer and Stewart, 1977; Zandt, 1978; Iyer, 1979], Long Valley, California [Steeles and Iyer, 1976b], and Kilauea Volcano, Hawaii [Ellsworth and Koyanagi, 1977].

In this paper the three-dimensional velocity structure under the Coso Range geothermal area, southern California, is studied using the *P* phases of steeply incident teleseismic waves. Lateral variations in velocity are estimated from observed *P* wave delays both by simple modeling employing ray tracing (the forward problem) and by the single-step three-dimensional inversion technique described by Aki et al. [1977].

The data are derived from approximately 5-km-wavelength plane compressional waves recorded by an array of sensors ~25 km in diameter with station spacing approximately 5 km in the center of the array (Figure 1 and Table 1). The array covers the area of Pleistocene rhyolite domes below the central portion of which the existence of a centralized magma reservoir has been proposed [Bacon et al., 1980]. *P* wave velocity in the proposed reservoir is expected to be abnormally low if magma or partially molten rock is currently present. With the collected suite of teleseismic data the three-dimensional inversion technique allows modeling of velocity structure in the volume beneath the array to a depth of 25 km with lateral resolution of approximately 5 km.

## DESCRIPTION OF THE SEISMIC ARRAYS

The U.S. Geological Survey (USGS) installed a network of 16 telemetered seismic stations in the Coso Range during September 1975 (Figure 1 and Table 1). This network was designed to provide continuous data for microearthquake studies and for teleseismic *P* wave delay studies. The network consisted of vertical geophones, seismic amplifiers, and telem-

etry electronics conforming to the USGS central California network station specifications [Bufe et al., 1975]. Recording was done at Menlo Park, California, on a photographic film recorder. One hundred thirty-seven teleseismic events, suitable for *P* wave delay study, were recorded from September 28, 1975, to May 9, 1976 (Table 2), and are referred to as the permanent network data.

During May and June 1977 the U.S. Geological Survey operated a portable seismic array (Centipede) in the Coso region (Figure 1 and Table 1). The array consisted of 25 1.0-Hz vertical geophones. For additional stations (RCW, HWS, CBH, and VPE) belonging to the Coso permanent network were recorded by both systems. Seismic amplifiers and telemetry electronics of the Centipede stations are described by Reasenbergs [1977]. The instrument response of the Centipede stations is nominally the same as that of the Coso permanent network. Recording was on a portable magnetic tape recorder located at Sugarloaf Mountain. Seven teleseismic events recorded during the 9 weeks of network operations were suitable for analysis and are referred to as the Centipede network data (Table 3).

## METHOD OF DATA REDUCTION AND ANALYSIS

The data considered in this study are the variations in the relative travel time between individual sensors as observed for steeply incident teleseismic *P* waves. These data are formed by differencing a predicted phase arrival time with arrival times measured on the seismograms. Relative travel time data for a dense, small-aperture network like the Coso array in essence map the cumulative phase distortion of the wave front created by its passage through heterogeneous structure encountered along its travel path. It is these relative delays in the wave front that are of central interest here and will be used to infer the presence (or absence) of lateral variations in velocity beneath the seismograph array.

*Measurements of P wave travel times.* Travel times for the permanent network data were determined visually by measuring the arrival time of the first peak and/or zero crossing of the *P* wave train. The uncertainty in timing clear arrivals is estimated to be better than 0.05 s. Arrival times for the Centipede network data were obtained by phase correlation of the *P* wavelets, as suggested by Press and Biehler [1964]. These measurements were made using an interactive computer program described by Reasenbergs [1978]. The principal advan-

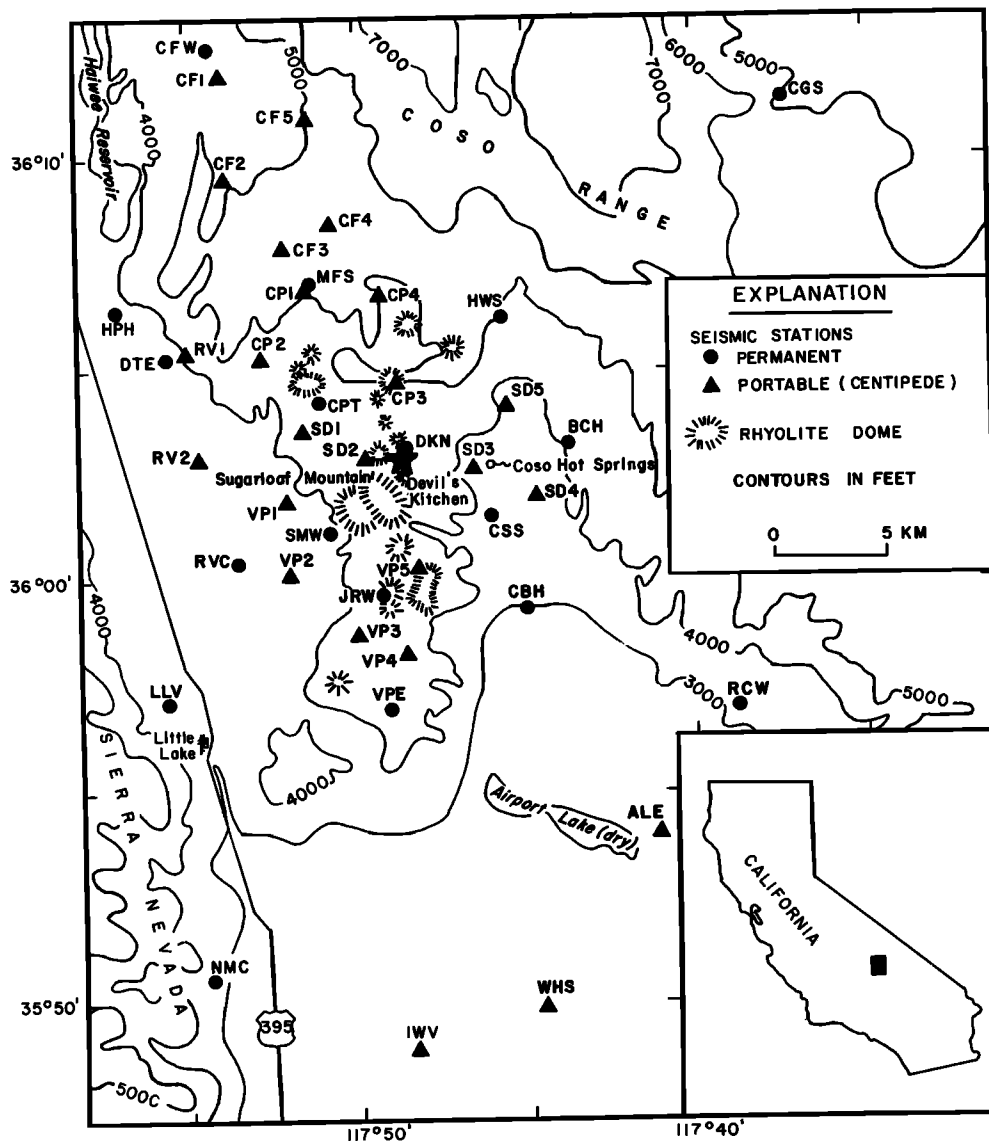


Fig. 1. Map of the Coso geothermal area showing generalized topography, locations of rhyolite domes, and seismograph stations used in this study. Central intersection of block element boundaries used in inversion models (Table 4) is indicated by star in Figures 1-6. Elevation contours in feet; (1 ft = 0.30 m).

tages of computer-assisted picking over hand processing include a dramatic improvement of timing accuracy, the ability to measure weak arrivals normally considered too small for visual analysis, and a measurement free from operator bias. The precision of the phase correlation time for each seismogram was estimated using the procedure described in the appendix. Timing uncertainties for the seven events analyzed averaged about 0.01 s.

**Relative travel time residuals.** Measured arrival times  $t_0$  were cast into a form suitable for further analysis by referencing them to a theoretical arrival time  $t_p$ , forming a travel time residual  $r = t_0 - t_p$ . To reduce the dependence of the residual data on total travel time and source origin time, relative residuals were formed by removing the mean value of the residuals from the individual reading on an event-by-event basis. The resulting zero mean relative residuals therefore do not depend on a reference station and depend only weakly on the specific set of stations with observations.

Two different methods were used to construct the reference travel time  $t_p$ . Source coordinates reported in the Preliminary

Determination of Epicenters bulletins of the U.S. Geological Survey were used together with the Herrin [1968] travel time table to construct one set of residuals. The second set was prepared for only the Centipede network data by determining the best fitting plane wave to the  $t_0$  readings for each event, using the method of least squares. The residual to this fitting of the data is the desired relative residual.

The relative residual patterns produced by these two methods are in close agreement (Figure 3). The insensitivity of the relative residuals to the particular reference wave front chosen demonstrates that the residual patterns observed reflect physical structure and are not merely artifacts of the analysis method. In the inversion analysis that follows, the Herrin residuals alone are used.

#### EVIDENCE FOR HETEROGENEOUS STRUCTURE FROM INSPECTION OF RELATIVE RESIDUAL DATA

The 137 events recorded by the permanent network were grouped according to source azimuths, and the mean relative residuals for each azimuthal quadrant were plotted and hand-

contoured (Figure 2). Inspection of the delay patterns reveals areas of excess delay in the southeast corner of the network for events from the northwest (Figure 2a) and in the northwest corner of the network for events from the southeast (Figure 2d). The large width of the source regions contributing to these residual patterns tends to defocus any heterogeneous structure that might be present, and only the most general pattern is revealed. Indeed, the azimuthally widest source region procedures the smoothest pattern (Figure 2b). Greater resolution of structural heterogeneity is obtained by observing waves from a smaller source region with a denser array. In

Figure 3, contoured relative residuals from the Centipede array are shown for two limited data sets, each representing a single source direction. A similar but more detailed delay pattern is revealed; the wave fronts from the northwest and southeast are delayed at stations southeast and northwest of the center of the array, respectively.

The azimuthal migration of accumulated wave front delays can most easily be explained as a 'delay shadow' produced by a low-velocity zone at depth. Accordingly, a crude model of a possible low-velocity zone can be made by ray tracing. Assuming a uniform upper crustal velocity of 6.0 km/s, a sphere

TABLE 1. Station Coordinates

Note	Station	Position	Elevation, m	Surface Element Solution
1	CF5	36°10.95'N, 117°51.35'W	1500	0.58
1	CF1	36°11.84'N, 117°53.78'W	1366	-0.36
1	CF4	36°8.46'N, 117°50.67'W	1646	1.92
1	CF3	36°7.89'N, 117°54.22'W	1527	2.13
1	CF2	36°9.48'N, 117°53.75'W	1451	2.93
1	CP1	36°6.94'N, 117°51.34'W	1518	1.50
1	CP2	36°5.27'N, 117°52.70'W	1320	1.12
1	CP4	36°6.78'N, 117°49.18'W	1522	-0.78
1	CP3	36°4.70'N, 117°48.74'W	1562	0.54
1	SD2	36°2.92'N, 117°49.68'W	1268	-2.07
1	SD1	36°3.54'N, 117°51.51'W	1268	-1.71
1	SD4	36°2.05'N, 117°44.70'W	1097	-4.64
1	SD5	36°4.19'N, 117°45.57'W	1170	-1.50
1	SD3	36°2.69'N, 117°46.55'W	1134	-4.42
1	VP1	36°1.87'N, 117°51.99'W	1622	6.04
1	VP2	36°0.15'N, 117°51.95'W	1094	1.55
1	VP5	36°0.34'N, 117°48.16'W	1268	1.27
1	VP3	35°58.71'N, 117°49.88'W	1513	4.23
1	VP4	35°58.27'N, 117°48.51'W	1501	1.55
1	RV2	36°2.86'N, 117°54.56'W	1059	0.34
1	RV1	36°5.37'N, 117°54.96'W	1195	-1.37
1	CLR	35°42.76'N, 117°35.80'W	663	-4.75
1	ALE	35°54.04'N, 117°41.16'W	732	
2	RCW	35°57.04'N, 117°38.86'W	945	
1	WHS	35°49.88'N, 117°44.60'W	707	
2	HWS	35°06.30'N, 117°45.67'W	1448	7.81
3	CBH	35°59.38'N, 117°45.01'W	884	-10.51
2	VPE	35°56.98'N, 117°49.02'W	1463	4.45
1	IWV	35°48.85'N, 117°48.31'W	706	
2	NMC	35°50.57'N, 117°54.29'W	951	
3	MFS	36°7.03'N, 117°51.30'W	1524	2.76
4	JRW	35°59.70'N, 117°49.20'W	1387	5.54
3	SMW	36°1.17'N, 117°50.72'W	1113	0.36
4	DKN	36°3.13'N, 117°48.56'W	1341	-1.25
4	RVC	36°0.47'N, 117°53.42'W	1066	2.49
5	CPT	36°4.26'N, 117°51.01'W	1494	-1.84
8	DTE	36°5.82'N, 117°55.52'W	1143	-4.23
6	HPH	36°6.38'N, 117°57.00'W	1160	
1	CGS	36°11.41'N, 117°37.39'W	1676	5.99
5	CSS	36°1.58'N, 117°46.01'W	1143	-1.64
7	BCH	36°3.28'N, 117°43.74'W	1265	-3.24
5	CFW	36°12.50'N, 117°54.23'W	1372	-1.74
9	LRS	36°12.38'N, 117°59.93'W	1326	
8	LLV	35°57.13'N, 117°57.13'W	975	
10	CLC	35°49.00'N, 117°35.80'W	766	-2.54

<sup>1</sup> USGS Centipede system portable station; operated May–June 1977.

<sup>2</sup> USGS permanent network station; operated September 1975 to present (July 1978).

<sup>3</sup> USGS permanent network station; operated September 1975 to October 1977.

<sup>4</sup> USGS permanent network station; operated September 1975 to March 1977 and July–October 1977.

<sup>5</sup> USGS permanent network station; operated September 1975 to March 1977 and July 1977 to present (July 1978).

<sup>6</sup> USGS permanent network station; operated March 1976 to October 1977.

<sup>7</sup> USGS permanent network station; operated September 1975 to March 1977.

<sup>8</sup> USGS permanent network station; operated September 1975 to March 1976.

<sup>9</sup> USGS permanent network station; operated April–June 1977.

<sup>10</sup> California Institute of Technology seismograph station.

TABLE 2. Teleseismic Events, Permanent Network

Date	Origin Time, UT	Depth, km	Distance, deg	Azimuth, deg	$dT/d\Delta$ , s/deg
Sept. 28, 1975	0356:5.2	612.	78	238	5.3
Sept. 28, 1975	0913:58.3	500.	84	234	4.9
Sept. 30, 1975	0350:59.2	135.	61	129	6.8
Sept. 30, 1975	0828:12.2	33.	46	310	7.9
Oct. 1, 1975	1801:6.4	556.	85	234	4.8
Oct. 2, 1975	1106:46.5	75.	70	310	6.1
Oct. 2, 1975	1840:49.2	617.	79	238	5.3
Oct. 3, 1975	0953:22.6	13.	49	310	7.6
Oct. 3, 1975	1416:17.1	56.	89	257	4.7
Oct. 6, 1975	0952:16.8	492.	85	233	4.9
Oct. 7, 1975	0637:57.4	115.	87	248	4.8
Oct. 10, 1975	0550:17.3	575.	82	236	5.0
Oct. 10, 1975	1313:9.4	96.	77	135	5.6
Oct. 11, 1975	1435:15.0	9.	81	231	5.3
Oct. 11, 1975	1441:13.6	63.	22	137	10.1
Oct. 11, 1975	1455:0.3	33.	81	231	5.3
Oct. 13, 1975	2051:20.2	35.	81	231	5.3
Oct. 13, 1975	2212:37.5	33.	81	231	5.3
Oct. 16, 1975	0337:42.5	33.	81	231	5.3
Oct. 17, 1975	0159:30.2	540.	84	236	4.9
Oct. 17, 1975	1612:30.4	267.	77	235	5.5
Oct. 17, 1975	1939:12.5	33.	30	326	8.9
Oct. 18, 1975	0859:56.3	1.	73	3	5.9
Oct. 19, 1975	0743:23.6	590.	78	238	5.3
Oct. 19, 1975	1218:5.9	83.	73	132	5.9
Oct. 19, 1975	1449:56.1	320.	85	293	4.9
Oct. 20, 1975	2225:29.0	33.	77	238	5.7
Oct. 21, 1975	1159:57.3	1.	71	2	6.1
Oct. 22, 1975	0510:42.9	429.	65	314	6.4
Oct. 23, 1975	0450:43.1	38.	72	236	6.0
Oct. 23, 1975	2318:7.8	52.	92	263	4.6
Oct. 28, 1975	1230:0.2	19.	2	42	14.5
Oct. 28, 1975	1445:58.7	28.	89	285	4.7
Oct. 29, 1975	0454:0.7	35.	25	134	9.4
Oct. 29, 1975	0631:4.1	214.	75	235	5.7
Oct. 30, 1975	0141:31.5	59.	73	310	5.9
Nov. 1, 1975	0117:33.9	113.	88	285	4.8
Nov. 1, 1975	0614:55.5	424.	78	237	5.4
Nov. 1, 1975	1840:30.7	33.	74	234	5.8
Nov. 3, 1975	0522:10.4	33.	82	156	5.2
Nov. 4, 1975	1205:56.9	24.	53	315	7.3
Nov. 4, 1975	1241:10.0	33.	56	323	7.2
Nov. 5, 1975	0158:54.4	44.	48	118	7.7
Nov. 5, 1975	0517:37.1	33.	81	88	5.3
Nov. 5, 1975	1704:21.4	12.	29	124	8.9
Nov. 6, 1975	0106:42.1	61.	48	311	7.7
Nov. 8, 1975	1100:24.5	77.	85	253	4.9
Nov. 8, 1975	1224:36.7	574.	81	236	5.1
Nov. 9, 1975	1530:5.9	77.	74	132	5.9
Nov. 9, 1975	1755:29.1	82.	84	288	5.0
Nov. 10, 1975	0443:8.2	33.	81	158	5.3
Nov. 11, 1975	0425:32.3	355.	69	313	6.1
Nov. 11, 1975	0854:18.5	31.	72	309	6.0
Nov. 13, 1975	0254:1.2	33.	36	315	8.6
Nov. 13, 1975	1548:46.0	69.	61	313	6.8
Nov. 14, 1975	0938:31.8	594.	78	238	5.3
Nov. 15, 1975	1528:30.4	33.	22	138	10.2
Nov. 16, 1975	2123:5.1	18.	69	311	6.3
Nov. 17, 1975	0645:47.2	122.	81	140	5.3
Nov. 19, 1975	0446:10.9	26.	58	9	7.0
Nov. 19, 1975	1106:27.5	62.	57	316	7.1
Nov. 22, 1975	0850:9.1	54.	42	310	8.1
Nov. 23, 1975	0939:28.1	227.	88	245	4.7
Nov. 23, 1975	2302:7.4	161.	75	310	5.7
Nov. 24, 1975	0334:49.0	33.	75	233	5.8
Nov. 24, 1975	2140:30.8	302.	84	232	4.9
Nov. 25, 1975	1415:2.1	593.	81	236	5.1
Nor. 27, 1975	0825:56.3	33.	66	53	6.5
Nov. 27, 1975	1042:47.7	33.	76	235	5.7
Nov. 27, 1975	2011:37.7	317.	85	293	4.9
Nov. 30, 1975	0531:25.7	99.	44	311	8.0
Dec. 2, 1975	2018:46.4	137.	85	288	5.0
Dec. 3, 1975	0546:40.1	246.	85	252	4.9
Dec. 5, 1975	1754:39.6	45.	36	315	8.6

TABLE 2. (continued)

Date	Origin Time, UT	Depth, km	Distance, deg	Azimuth, deg	$dT/d\Delta$ , s/deg
Dec. 5, 1975	2014:18.2	68.	70	310	6.2
Dec. 6, 1975	2247:30.4	82.	75	134	5.8
Dec. 8, 1975	1855:40.6	54.	58	315	7.0
Dec. 9, 1975	0914:40.6	33.	73	236	5.9
Dec. 9, 1975	1334:4.1	635.	79	238	5.3
Dec. 11, 1975	0516:8.4	50.	86	253	4.9
Dec. 15, 1975	0424:28.8	200.	86	230	4.8
Dec. 19, 1975	0214:29.6	33.	87	253	4.8
Dec. 21, 1975	1054:17.7	554.	63	316	6.5
Dec. 21, 1975	1751:4.8	109.	83	291	5.1
Dec. 26, 1975	1556:38.7	33.	73	235	5.9
Jan. 1, 1976	0129:39.6	59.	85	230	4.9
Jan. 1, 1976	0221:10.6	50.	86	229	4.9
Jan. 1, 1976	1903:29.4	63.	86	229	4.9
Jan. 2, 1976	0410:14.8	64.	86	229	4.9
Jan. 2, 1976	0950:20.6	50.	86	285	4.9
Jan. 5, 1976	0231:36.3	95.	64	132	6.6
Jan. 6, 1976	1928:36.5	33.	56	317	7.2
Jan. 6, 1976	2354:22.2	76.	70	131	6.1
Jan. 7, 1976	1832:10.2	58.	92	305	4.6
Jan. 9, 1976	2354:35.6	168.	87	248	4.8
Jan. 12, 1976	0656:31.3	634.	81	236	5.0
Jan. 13, 1976	1329:19.5	33.	62	27	6.7
Jan. 14, 1976	1530:4.5	407.	82	304	5.1
Jan. 16, 1976	1046:15.1	487.	85	234	4.9
Jan. 16, 1976	1054:28.0	33.	90	244	4.7
Jan. 16, 1976	2332:9.4	51.	86	229	4.9
Jan. 18, 1976	0446:24.4	33.	63	9	6.6
Jan. 20, 1976	1722:15.2	544.	80	236	5.2
Jan. 21, 1976	0601:50.9	33.	61	335	6.7
Jan. 21, 1976	1022:49.0	50.	68	309	6.3
Jan. 22, 1976	1828:58.3	56.	92	263	4.6
Jan. 27, 1976	2328:20.9	394.	82	303	5.1
Feb. 5, 1976	0953:11.7	98.	74	133	5.8
Feb. 5, 1976	1713:12.5	78.	71	309	6.1
Feb. 9, 1976	2129:57.1	48.	17	143	12.3
Feb. 14, 1976	1050:22.2	548.	84	298	4.9
Feb. 18, 1976	0800:58.6	39.	45	310	7.9
Feb. 28, 1976	1627:9.0	9.	85	148	4.9
March 4, 1976	0250:0.5	90.	87	249	4.8
March 8, 1976	2006:33.3	72.	86	252	4.9
March 13, 1976	2144:41.3	165.	50	114	7.5
March 25, 1976	2216:10.6	33.	43	138	8.1
March 28, 1976	2019:45.6	33.	63	66	6.6
April 4, 1976	2238:29.5	332.	65	314	6.4
April 7, 1976	0038:28.0	33.	78	233	5.6
April 10, 1976	1712:9.2	560.	78	238	5.4
April 11, 1976	1301:49.6	545.	80	316	5.2
April 11, 1976	1303:35.7	529.	80	316	5.2
April 12, 1976	0441:51.4	38.	40	311	8.3
April 20, 1976	1427:20.6	30.	28	336	8.9
April 24, 1976	1105:15.7	94.	78	138	5.5
April 27, 1976	1814:19.0	33.	25	136	9.3
May 4, 1976	0440:46.5	33.	77	89	5.6
May 5, 1976	1637:9.1	73.	74	133	5.8
May 9, 1976	2044:44.7	34.	92	263	4.6

of radius 5 km buried beneath Devil's Kitchen at 15-km depth, with velocity 5.4 km/s, could roughly account for the location and amplitude of the observed wave front delays. Such a model, however, is highly nonunique and serves mainly as a guide for formulation of the three-dimensional inversions presented below.

#### DETERMINATION OF A THREE-DIMENSIONAL VELOCITY STRUCTURE

Variations in the pattern of relative travel time delays described above present clear evidence for the presence of a low-velocity body imbedded in the crust beneath the Coso Range. A more complete modeling of these data using the three-di-

mensional inversion method of *Aki et al.* [1977] permits better definition of the body and measurement of the velocity contrast between it and the host rock. The method determines velocity perturbations within a limited volume underlying the seismograph array that minimize the variance of the observed travel time data through solution of a linearized system of equations.

*Methodology.* The modeling procedure used in this paper closely follows the single-step stochastic inverse solution described by *Aki et al.* [1977]. An initial model composed of homogeneous plane parallel layers is divided into a grid of right rectangular prisms. The inverse solution determines perturbations in velocity, with respect to the homogeneous start-

TABLE 3. Teleseismic Events, Centipede System Array

Event	Origin	Date 1977	Time, UT	Magnitude	Position	Distance, deg	Azimuth, deg	$dT/d\Delta$ , s/deg
1	Mariana Islands	May 24	1023:23.4	5.7	18.8°N, 145.4°E	85	289	4.91
2	Tonga	June 1	0857:30.9	5.2	21.2°S, 174.4°W	77	234	5.62
3	Chile/Bolivia	June 8	1325:15.6	5.4	22.1°S, 67.3°W	74	132	5.82
4	Honshu, Japan	June 8	1425:46.5	5.5	38.5°N, 141.5°E	76	308	5.66
5	North Atlantic Ridge	June 14	2352:58.6	5.3	16.6°N, 46.6°W	67	87	6.42
6	Mariana Islands	June 17	1445:09.4	5.6	19.0°N, 145.7°E	83	289	5.14
7	Tonga	June 22	1208:33.4	6.8	22.9°S, 175.9°W	80	233	5.40

Data are from the U.S. Geological Survey, National Earthquake Information Service. Preliminary Determination of Epicenters.

ing model, for each prism or model element. Ray paths along the reference wave front are traced through the block structure for each source-receiver pair, and the time spent by each ray in each block  $a_{ijk}$  is calculated for use in constructing normal equations relating fractional changes in block velocity  $m_k$  to the zero mean time residuals  $r_{ij}$ . The  $k$ th normal equation is

$$\sum_{ij} \left( a_{ijk} a_{ijl} - \frac{\bar{a}_{jk} \bar{a}_{jl}}{n_j^2} \right) m_k = - \sum_{ij} a_{ijk} \left( r_{ij} - \frac{\bar{r}_j}{n_j} \right) \quad (1)$$

where

$$\sum_i a_{ijk} = \bar{a}_{jk}$$

$$\sum_i r_{ij} = \bar{r}_j$$

Using matrix notation, (1) may be written as

$$\mathbf{G} \mathbf{m} = \mathbf{d} \quad (2)$$

where  $\mathbf{m}$  is a vector containing the yet unknown model perturbations  $m_k$ ,  $\mathbf{G}$  is a semidefinite matrix defined by the left side of (1), and  $\mathbf{d}$  is a vector defined by the right side of (1). The method of damped least squares [Levenberg, 1944; Marquardt, 1963] is used to solve (1), and the solution is given by

$$\hat{\mathbf{m}} = (\mathbf{G} + \theta^2 \mathbf{I})^{-1} \mathbf{d} \quad (3)$$

where  $\hat{\mathbf{m}}$  is the damped least squares solution,  $\theta^2$  is a positive constant (damping parameter), and  $\mathbf{I}$  is the identity matrix. For (3) the resolution matrix is given by

$$\mathbf{R} = (\mathbf{G} + \theta^2 \mathbf{I})^{-1} \mathbf{G} \quad (4)$$

and the covariance matrix by

$$\mathbf{C} = \sigma_d^2 (\mathbf{G} + \theta^2 \mathbf{I})^{-1} \mathbf{R} \quad (5)$$

when data errors are uncorrelated and have uniform variance  $\sigma_d^2$ . The standard error of the solution is given by

$$\Delta \hat{m}_i = C_{ii}^{1/2}$$

and is bounded by

$$\Delta \hat{m}_i \leq \frac{\sigma_d}{\theta} [R_{ii}(1 - R_{ii})]^{1/2} \leq \frac{\sigma_d}{2\theta} \quad (6)$$

[Ellsworth, 1977].

Selection of a value for the damping parameter  $\theta^2$  is motivated by Aki et al. [1977], who note that when

$$\theta^2 = \sigma_m^2 / \sigma_m^2 \quad (7)$$

(3) is a special case of the stochastic inverse of Franklin [1970], where  $\sigma_m$  is the root mean square fluctuation of the true model  $\mathbf{m}$ .

Solution of the linearized problem defined by (1) produces an image of lateral variations in velocity within each layer of the model. Absolute values of the velocity variations cannot be determined by the method [Aki et al., 1977]. Consequently, the solution within each layer is relative to an unknown mean. Fortunately, precise knowledge of the average vertical velocity profile is not required for the solution to have quantitative value, since the method can tolerate substantial errors in layer velocity. Iterative refinement of the solution would be desirable to help reveal the fine details of the velocity heterogeneities. However, numerical experiments reported by Ellsworth [1977] and Smith et al. [1978] show the linear theory to be remarkably accurate. All solutions reported below are single-step solutions.

*Application to Coso.* Successful imaging of laterally varying structure requires an appropriate distribution of both sources and receivers. Combination of data from the permanent and Centipede networks provides an ideal data set of azimuthally well distributed source events with travel times measured on a dense grid of receivers. These data allow a horizontal spacing of block model elements of 5 km (Table 4), which effectively defines the lateral resolution of the model. The height of the blocks is chosen to enhance the vertical resolution. Cubic blocks cannot be adequately resolved by the data owing to the steep angle of incidence of teleseismic  $P$  waves in the crust. Vertical resolution can be improved by selecting blocks that are taller than they are wide, since this improves the crossfire of ray paths through the model [Ellsworth and Koyanagi, 1977]. With the exception of the surface layer a height-to-width ratio of 1.5, for a layer thickness of 7.5 km, is used. This layer thickness provides useful vertical resolution in the crust.

Modeling of the surface layer requires special treatment because ray paths to individual stations do not mix with ray paths from neighboring stations above about 5-km depth. Without loss of generality, each station may then be assigned a unique first-layer element in place of a regular block format for this layer. Inversion of this three-layer model simultaneously determines average velocity variations for the block elements in layers 2 and 3 and surface element solutions (Table 1) for layer 1. Use of these surface element solutions (effectively, angle-of-incidence dependent station corrections) improves the numerical resolution of the underlying layers.

The layered earth model used in computing ray paths is a simplified version of the reversed refraction model determined by C. S. Weaver and A. Walter (unpublished manuscript, 1980). The most extreme mismatch between the approximate layered structure and the measured structure occurs for the model with the thickest first layer (Table 4). Numerical experiments show that the use of a velocity 25% too low for this layer does not significantly affect the inverse solution.

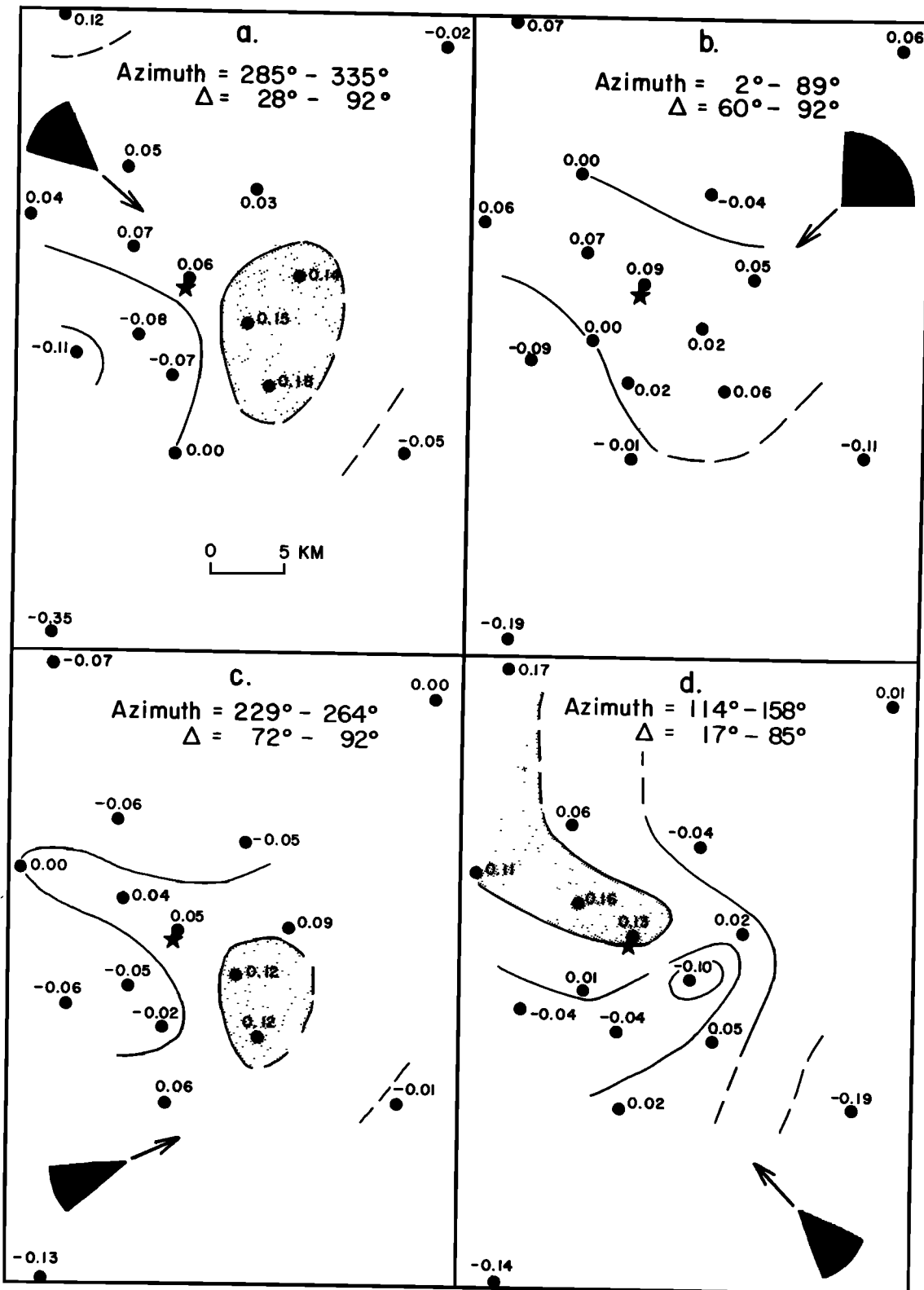


Fig. 2. Relative residual patterns derived from permanent network data. Shown for each station are mean Herrin residuals for events within four limited source regions indicated. Corresponding data uncertainties average 0.02 s. Contour interval, 0.1 s. Shading covers areas where delay exceeds 0.1 s.

Selection of the proper value for the damping parameter using (7) implies that  $\sigma_m$  be known. Since  $\sigma_m$  is initially unknown, a series of model experiments was conducted to establish a range of values for various choices of  $\theta$ . The results of three experiments (Table 5) indicate that the mean square

fluctuation of the crustal velocity beneath the array is about 2.5%, which suggests an ideal value of  $\theta^2 = 0.0004 \text{ s}^2/\%^2$ . Table 5 also illustrates the classical trade-off between resolution and modeling error discussed by *Backus and Gilbert* [1968]. The value of  $\theta^2 = 0.001 \text{ s}^2/\%^2$  adopted for the final



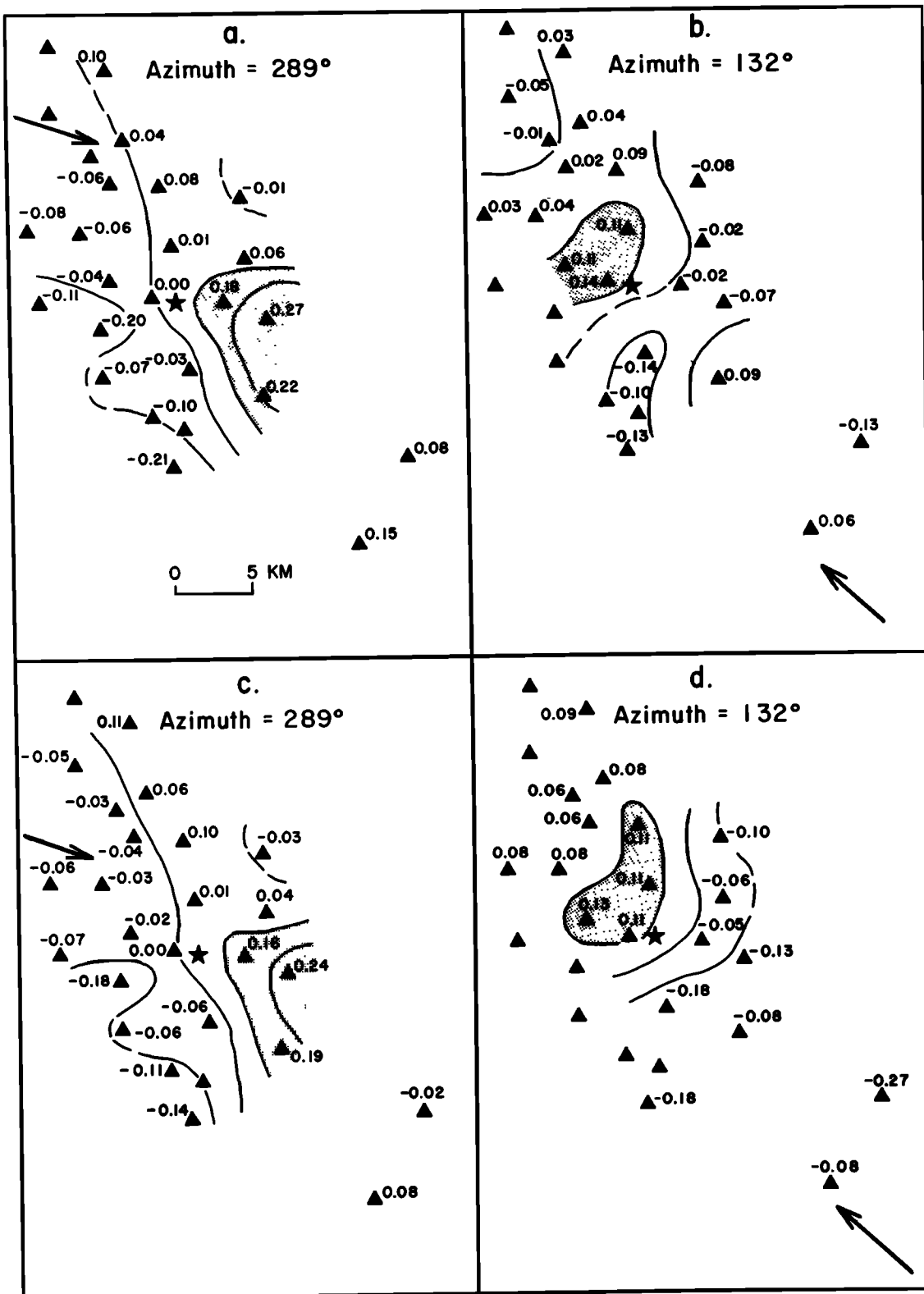


Fig. 3. Relative residual patterns for events from two particular source directions, recorded by Centipede array. Herrin residuals are shown in Figures 3c and 3d; least squares fitted plane wave residuals are shown in Figures 3a and 3b. Contour interval, 0.1 s.

models discussed below reflects our preference for improved control over standard errors at the expense of a slight degradation in spatial resolution. Explicit examples of model resolutions are discussed below. Using the adopted value for  $\theta^2$ , the depth to which modeling of laterally varying structure is

feasible was determined by the  $F$  test introduced by Aki *et al.* [1977] to be 20–25 km (a three-layer model).

*Resolution of three-dimensional structure.* Application of the inversion method to the Coso network determines a high-resolution image of lateral variations in velocity structure be-

TABLE 4. Layered Velocity Models Used for Inversion Studies

Layer	P Velocity, km/s	Thickness, km	Block Length, km
<i>Model A</i>			
1	4.5	2.5	5.0
2	6.0	7.5	5.0
3	6.0	7.5	5.0
<i>Model B</i>			
1	4.5	5.0	5.0
2	6.0	7.5	5.0
3	6.0	7.5	5.0
<i>Model C</i>			
1	4.5	7.5	5.0
2	6.0	7.5	5.0
3	6.0	7.5	5.0

neath the array (Figure 4). Rapid spatial variations in velocity with spatial wavelengths comparable to those seen in the wave front delay diagrams (Figures 2 and 3) are clearly present. The modeled structure also indicates the continuance of significant lateral heterogeneity through the deepest level modeled.

The inversion results for a single model depend strongly upon the precise configuration of the block elements because the solution for a block represents the average velocity fluctuation in its volume. The influence of the block boundaries on the solution can be reduced by forming the spatial average of two models with block gridwork displaced diagonally by one-half block length. These smoothed images of lateral heterogeneity are suitable for contouring (Figure 5). The position of the layer boundaries also exerts an artificial influence upon the model. Consequently, three different divisions of the crust into layers are presented (Table 4). Although these models are not independent, they illustrate the vertical transformation of the heterogeneity pattern.

The ensemble of models, considered individually, succeeds in explaining the travel time observations. For example, the solution depicted in Figure 4 reduces the variance of the residuals by 80%, leaving an unmodeled rms residual of 0.047 s. Although the models adequately explain the data using less than 10% of the available degrees of freedom, the uniqueness and significance of the models must be quantified before a reliable geologic interpretation can be made.

The inherent nonuniqueness of the inverse solution can be

studied in the framework of the linearized theory using the resolution matrix defined in (4). Each row of the resolution matrix acts as a filter on the true model  $\mathbf{m}$  which is unattainable, yielding the damped least squares solution  $\hat{\mathbf{m}}$ :

$$\hat{\mathbf{m}} = \mathbf{R} \mathbf{m}$$

For the model to be well resolved, the rows of  $\mathbf{R}$  should act as narrow filters, ideally delta functions, centered on the main diagonal of  $\mathbf{R}$ . The presence of nonzero elements off the main diagonal quantitatively expresses the smoothing of  $\hat{\mathbf{m}}$  created by the inverse operator in (3).

Study of the resolution matrix for the model in Figure 4 shows that blocks with  $R_{ii} \geq 0.6$  are adequately resolved by the data. These blocks characteristically underlie the region enclosed by the array. Peripheral blocks are poorly resolved in general and cannot be interpreted reliably. Rows of  $\mathbf{R}$  for selected blocks in layer 2 of Figure 4 appear in Figure 6. They show that the solution for these blocks is effectively decoupled from the surface layer. Smoothing between layers 2 and 3 is stronger. However, it is not judged to be strong enough to distort the broad-scale feature of the solution.

The significance of the perturbations determined for each block is estimated using the model standard error  $\Delta\hat{\mathbf{m}}$ . Accepting the residual variance as a measure of the true data variance, (6) gives an upper bound on  $\Delta\hat{\mathbf{m}}$  of 0.73%. Accordingly, we adopt 1.5% as a figure of merit in interpreting the significance of the models.

## DISCUSSION

*Low-velocity zone below the geothermal area.* The three-dimensional velocity structure beneath the array clearly reveals a central, intense low-velocity body in the middle crust (Figures 4 and 5). The structure has a width of approximately 5 km and becomes increasingly elongated in the N-S, or N25°E, direction with increasing depth. In each layer except the deepest the -1% velocity contour surrounds a single, simply connected zone. The maximum velocity contrast is between 5.6 and 8.4% and attains its maximum value between 10 and 17.5 km.

Although this ensemble of layered models demonstrates the general properties of the lateral inhomogeneity over the entire depth range modeled, the details of the velocity structure are best seen by considering the unsmoothed solutions for the middepth model (Figures 4 and 6). The solution for the surface layer (Figure 4a) shows below average velocity in the

TABLE 5. Effect of Varying Damping Parameter on Average Performance of Inversion Model

Damping Parameter $\theta^2, * s^2/\%^2$	Assumed $\sigma_m$ , %	rms Velocity Perturbation, %	Mean Standard Error, %
0.0003	3.0	2.48	0.87
0.0010	1.6	2.22	0.64
0.0030	1.0	1.95	0.44
Theoretical Maximum Standard Error, † %	Mean Resolution	Tr (R) ‡	Data Variance Reduction, %
1.61	0.83	75	82
0.89	0.69	57	80
0.52	0.51	40	76

Values are averages of central 16 elements from layers 2 and 3.

\* Standard error of data assumed to be 0.05 s.

† Upper bound on standard error given by (6) when  $R_{ii} = 0.5$ .

‡ Trace of resolution matrix for all 102 free parameters (number of singular vectors of  $\mathbf{G}$  that effectively contribute to the solution).

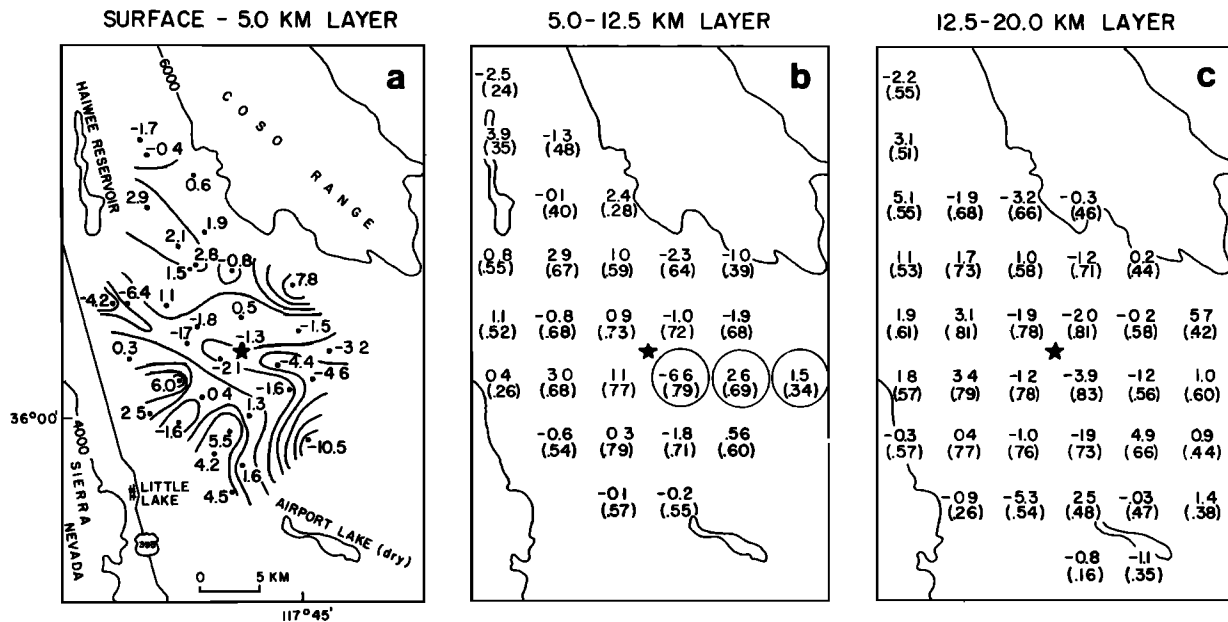


Fig. 4. Unsmoothed inversion solution  $\mu$  for model B (Table 4) showing percent velocity variations, with respect to homogeneous starting model, for model elements. (a) In the surface layer, each model element is associated with an individual seismograph station. Solutions for these elements are shown next to their corresponding stations and listed in Table 1. Areas of low surface velocity (negative solutions) correspond to locations of sedimentary basins (see text). (b, c) Percent velocity variations for blocks in layers 2 and 3 are shown at center of each solution element. In parentheses are the corresponding resolution diagonal elements  $R_{ii}$ . Circled values indicate blocks for which the full resolution matrices are shown in Figure 6.

sedimentary areas of Cactus Flat (6 km east of Haiwee Reservoir), Upper Cactus Flat (9 km north of Sugarloaf Mountain), Rose Valley (south of Haiwee Reservoir), and Coso Basin (7 km north of Airport Lake). A significant ridge of low velocity oriented approximately east-west crossing the central part of the array is possibly associated with the fractured and altered rock in that area (C. Bacon, personal communication, 1979). The solution for the surface is well decoupled from the layer below it (Figure 6) and is most easily explained in terms of low-velocity surficial deposits. Within the well-resolved ( $R_{ii} \geq 0.6$ ) portion of layer 2 (Figure 4b), significantly low velocity rock is contained in a circular region 5–10 km in diameter, centered 3 km southeast of the grid center. The lowest velocity in this layer is attributed to the block located beneath Devil's Kitchen. Within the well-resolved area of the deepest layer (Figure 4c) the low-velocity zone is elongate and oriented approximately N-S, with lateral dimension 10–25 km. The lowest-velocity block in this layer lies directly below the lowest-velocity block in layer 2. Because of some significant vertical smoothing of the solution between these two blocks (Figure 6), part of the solution for the layer 3 block could be due to the extremely low velocity value in layer 2.

**Partial melt.** The observed lateral variations in velocity could arise through several distinct physical mechanisms. Iyer and Stewart [1977] have presented a thorough review summarizing the effects upon compressional velocity of temperature, stress, density, mineral orientation, crack and pore properties, and fluid inclusions (including partial melt). We do not undertake to evaluate each of these possibilities for Coso but rather explore the implications of what we consider the most likely candidate, the presence of partial melt.

Partial melt has been proposed as a possible explanation for low-velocity zones inferred from  $P$  wave delays at Long Valley [Steeple and Iyer, 1976b], Yellowstone [Iyer and Stewart, 1977], and The Geysers [Iyer et al., 1979]. If partial melt is

present at Coso, it can sufficiently explain the 6–8% crustal velocity variations observed. Murase and McBirney [1973] have demonstrated that partial melting of basalt, occurring at approximately 800°C, is capable of reducing compressional velocity by more than this amount.

It is well established that molten rock has recently been present at (or near) the surface of the geothermal area. The thermal history at Coso over the past 4 m.y., as revealed by the geology, is characterized by two distinct periods of volcanism, the more recent of which included the emplacement of the rhyolite domes (Figure 1) between 0.3 and 0.04 m.y. ago [Duffield et al., 1980]. Estimates of the present physical state of the crust at Coso are suggested by local and regional seismic and teleseismic evidence. The apparent confinement of the local and regional seismicity to the upper 8–10 km [Walter and Weaver, 1980] implies ductile behavior below this depth; this implied ductile behavior suggests that some combination of high temperature, low strain rate, and low fluid pressure exists on a regional scale [Verhoogen et al., 1970; R. Stewart, personal communication, 1979]. This depth corresponds closely to the depth of the low-velocity zone revealed by the teleseismic arrivals. The lateral location of the low-velocity zone coincides with the area of highest heat flow measurements [Combs, 1980], hot springs, fumaroles, and most recent volcanism [Duffield et al., 1980]. To explain the presence of a localized low-velocity zone within a regionally hot and ductile crust requires an additional mechanism. Local presence of partial melt is suggested by the thermal and geologic evidence and is consistent with seismic evidence.

The direction of the least compressive stress in the Coso geothermal area is determined by fault-plane solutions [Weaver and Hill, 1979; Walter and Weaver, 1980] and by surface faulting [Bacon et al., 1980] to be oriented approximately east-west. Dynamic models for Coso that include partial melting and are consistent with the local stress regime have been

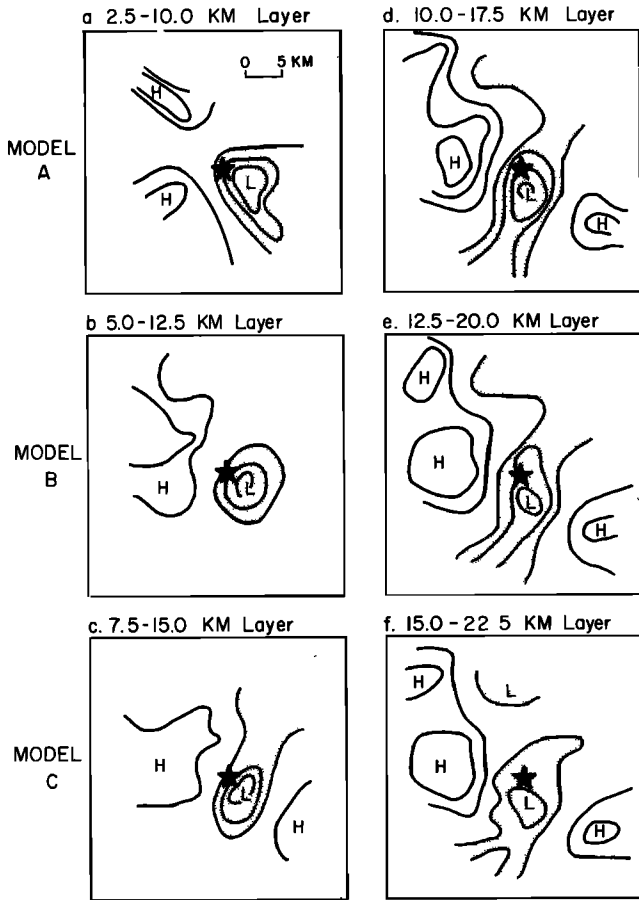


Fig. 5. Smoothed inversion solutions for the second and third layers of the three-block layer divisions modeled (Table 4). Velocity variation contours at 1% intervals (0% contour omitted) are shown. High-velocity and low-velocity zones are indicated by H and L, respectively. Shaded areas indicate compressional wave velocity 1% or more slower than average layer velocity.

proposed by *Weaver and Hill* [1979], *Bacon et al.* [1980], and *Duffield et al.* [1980]. They suggest that the thermal system which resulted in the deposit of rhyolite at the surface originated with the passive intrusion of basaltic magma into extensional fractures, producing north-south aligned, nearly vertical dikes.

Similar models have been proposed by *Lachenbruch et al.* [1976] for Long Valley and by *Thompson and Burke* [1974] as a general mechanism for the Basin and Range province. This model is consistent with the evidence from the teleseismic arrivals at Coso. The close coincidence of the surface projection of the low-velocity zone with the area of rhyolite domes (Figure 1) suggests its identification with the proposed rhyolite reservoir and intrusive dike beneath it. The depth and dimensions of the proposed dike-melt system agree with those estimated for the teleseismic low-velocity zone.

CONCLUSIONS

The pattern of relative residuals derived from teleseismic arrivals recorded at the Coso geothermal area reveals an area of excess delay that migrates with changing source azimuth. This migration can most easily be explained as a delay shadow produced by a low-velocity zone at depth. A simple, nonunique model consisting of a 5-km-radius sphere with 10% lowered velocity buried 15 km below Devil's Kitchen can roughly account for the amplitude and location of the observed days.

Inversion of the relative residual data for three-dimensional velocity structure determines a high-resolution image of lateral variations in velocity beneath the seismic array. The modeled three-dimensional velocity structure successfully explains the travel time observations by reducing the variance of the residuals approximately 80%.

An intense, low-velocity body is revealed between 5- and 20-km depth under Devil's Kitchen, approximately 5 km wide on top, becoming increasingly elongated in the N-S, or

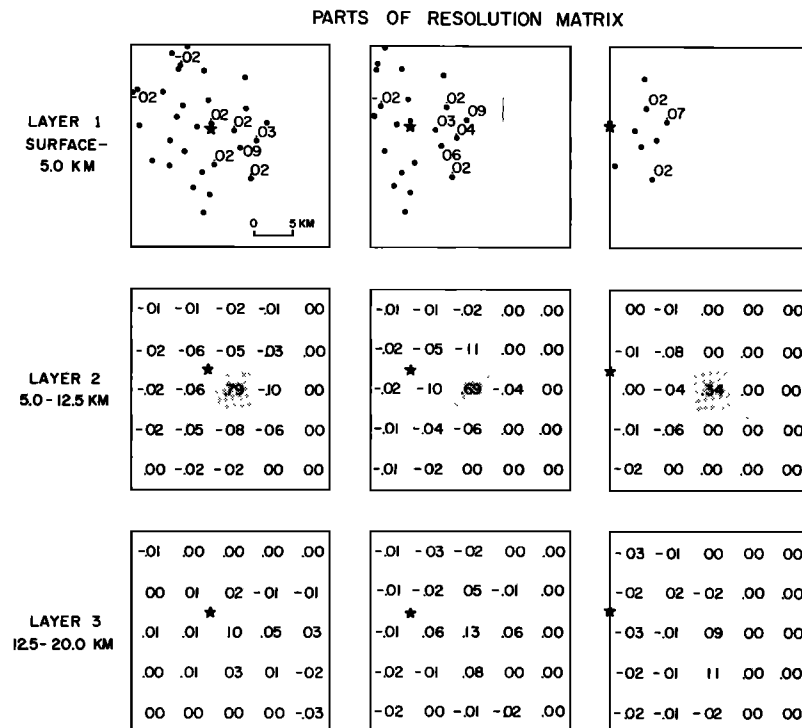


Fig. 6. Selected elements of the resolution matrix for circled block elements in Figure 4. Diagonal element is shaded. Off-diagonal elements in layer 1 smaller than 0.02 are denoted by a dot at the station location.

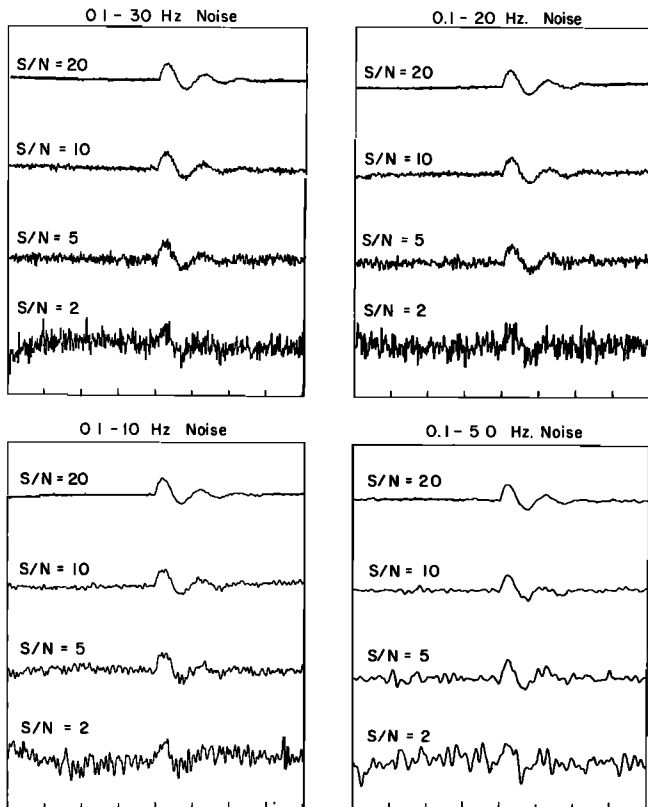


Fig. 7. Examples of synthetic teleseismic  $P$  pulses selected from ensembles of traces corresponding to the signal-to-noise ratio and noise bandwidth shown. Tick marks are seconds. Correlation window includes the first 2 s of the pulse.

TABLE 6. Expected Timing Errors Determined From Synthetic Seismograms

$S/N$	$S/N$			
	20	10	5	2
<i>0.1-to 30-Hz Noise</i>				
20	3	4	8	23
10		7	10	21
5			17	33
2				169
<i>0.1-to 20-Hz Noise</i>				
20	8	5	11	26
10		9	10	25
5			16	34
2				144
<i>0.1-to 10-Hz Noise</i>				
20	9	8	16	115
10		14	14	114
5			24	71
2				255
<i>0.1-to 5-Hz Noise</i>				
20	6	9	21	202
10		10	28	173
5			52	203
2				509

Values are expected timing errors (in milliseconds) determined from an ensemble of 71 cross correlations involving traces with signal-to-noise ( $S/N$ ) ratio and noise spectrum indicated. For example, for pairs of traces having  $S/N$  ratios of 10 and 2, each with noise bandwidth of 0.1–30 Hz, the expected timing error is 21 ms. Length of the correlated trace segments is 2 s. Figure 7 shows examples of synthetic seismograms from each ensemble.

$N25^{\circ}E$ , direction with increasing depth. Maximum velocity contrast of 8.4% is attained between 10- and 17.5-km depth.

A plausible explanation for the existence of the low-velocity zone can be made by hypothesizing the presence of partial melt in the middle crust. This hypothesis is consistent with the geologic, thermal, and local seismic evidence. A mechanism for partial melting, proposed by Duffield *et al.* [1980] and by Bacon *et al.* [1980] postulates the presence of a N-S aligned intrusive dike-melt system under the rhyolite field that coincides with the location and orientation of the low-velocity zone. Agreement between the velocity model obtained from teleseismic arrivals and this dynamic geologic model lends support to both models.

#### APPENDIX: ESTIMATION OF PRECISION OF TIMING BY PHASE CORRELATION

The precision with which the  $P$  phases of seismograms from an array can be relatively timed by cross correlation varies with the signal-to-noise ratio and noise spectral content of the seismograms and with the degree to which the  $P$  pulse shape remains constant across the array. Because the Centipede array size is small (Figure 1), the  $P$  pulse shape is similar from station to station, so that the correlation technique is appropriate for the Centipede data. Estimates of timing precision for the Centipede data were obtained statistically by cross-correlating pairs of traces from ensembles of synthetic seismograms. Each synthetic seismogram was composed of a simple function resembling a teleseismic  $P$  pulse to which was added a unique sample of stationary, band-limited Gaussian noise. Each ensemble, consisting of 71 such traces, was characterized by a particular signal-to-noise ratio and noise spectrum (Figure 7). Since the synthetic pulse function is identical in all traces (only the noise portion varied), the expected time lag between any two synthetic traces is zero. The cross correlation of all traces in an ensemble results in a distribution of estimates of lag between traces. The mean of the distribution tends to zero. The standard deviation provides a chi squared estimator of the expected error (due to noise) in the lag determined by the cross correlation. For each pair of ensembles corresponding to a particular combination of  $S/N$  ratio and noise spectrum the expected error estimate is shown in Table 6. These error estimates were applied to the recorded data by visually matching each recorded trace with the synthetic trace most closely resembling it (Figure 7) and adopting the corresponding synthetic trace error estimate. The resulting estimates of timing precision from the Centipede data range from 4 to 8 ms for the best events to 32 ms for the weakest arrivals.

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